

Ion-Molecule Chemistry of  $H_nC_3O^+$ ,  $C_3O_2^+$ , and  $C_3O^+$ Simon Petrie,<sup>†</sup> Ryan P. A. Bettens,<sup>‡</sup> Colin G. Freeman,<sup>\*</sup> and Murray J. McEwan

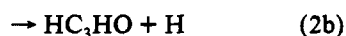
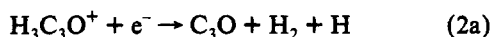
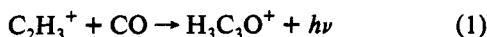
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Rate coefficients and product distributions have been measured using a selected ion flow tube for reactions of the ions  $H_nC_3O^+$  ( $n = 0-3$ ) and  $C_3O_2^+$  with the neutral species  $H_2$ ,  $CO$ ,  $N_2$ ,  $O_2$ ,  $CH_4$ ,  $NH_3$ ,  $H_2O$ ,  $C_2H_2$ ,  $HCN$ , and  $CH_3CCH$  at 300 K. The structures of the ions are discussed, and a survey of relevant enthalpies of formation is presented.

## Introduction

Tricarbon monoxide,  $CCCO$ , and propynal,  $HCCCHO$ , are among the growing list of molecules that have been observed in interstellar clouds.<sup>1,2</sup> Although the  $H_nC_3O^+$  ions have not been observed in the interstellar medium, they are believed to have a part in the production of  $C_3O$  and propynal. One of the mechanisms suggested for the production of these species in the interstellar medium is dissociative recombination of the  $H_3C_3O^+$  ion, which can be produced as a result of the radiatively-stabilized association reaction between  $C_2H_3^+$  and  $CO$ .<sup>3,4</sup> Other unsaturated

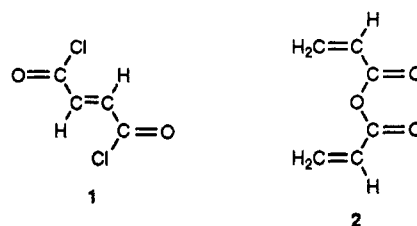


organo-oxygen molecular ions can be formed by the association of unsaturated hydrocarbon ions  $C_xH_y^+$  ( $x = 2$  or  $3$ ) with  $CO$  and perhaps also by reaction of these ions with atomic oxygen,<sup>4,5</sup> molecular oxygen,  $H_2O$ , and  $CO_2$ .<sup>6</sup>

In the interstellar context it is important to establish the behavior of the  $H_nC_mO^+$  ions, for if they are to play a role in the synthesis of propynal and tricarbon monoxide, then their lifetimes must be sufficiently long for dissociative recombination processes to present a significant loss mechanism. Their chemical reactivity therefore needs to be established. Few studies have been published on the reactivity of  $H_nC_mO^+$  ions, and the present work was undertaken to examine the reactivity of a range of these ions, with  $m = 3$ , with selected neutral reactants.

## Experimental Section

The selected ion flow tube (SIFT) apparatus and the procedures used to measure the rate coefficients and branching ratios reported here have been described previously.<sup>7</sup> In brief, ions produced by electron impact on a suitable reagent are mass-selected and injected into a flow tube. Neutral reactants are introduced into the flow tube at fixed reaction ports, and the ensuing reactions are monitored by a second quadrupole mass spectrometer located behind a small sampling orifice at the downstream end of the flow tube. Electron impact (typically 30 eV) on appropriate source gases produced the required  $H_nC_3O^+$  ions:  $HC_3O^+$  and  $H_2C_3O^+$  from fumaroyl dichloride (1) and  $H_3C_3O^+$  from acrylic anhydride (2).  $C_3O_2^+$  and  $C_3O^+$  were both formed in the ion source by



electron impact on carbon suboxide,  $C_3O_2$ . All reagents, except for acrylic anhydride and carbon suboxide, were obtained commercially and were generally purified by additional vacuum distillation. Acrylic anhydride was prepared from acrylic acid which was treated with triethylamine followed by phosgene, before purification by filtration and distillation.<sup>8</sup> Carbon suboxide was generated by the action of  $P_2O_5$  on malonic acid.<sup>9,10</sup> All measurements were taken at room temperature ( $295 \pm 5$  K) and at a helium bath gas pressure of 0.30 Torr.

**Identity of Ions Formed by Electron Impact.** The  $H_nC_3O^+$  ions produced in the manner described above are all assumed to exist as the lowest-energy isomers, and the relevant thermodynamic data for these species, along with some of their isomeric counterparts, are summarized in Table I. No isomer of  $HC_3O^+$ , other than  $HC\equiv C-C=O^+$ , has been identified, and we assign the  $HC_3O^+$  ion produced from fumaroyl dichloride to this structure. The  $H_2C_3O^+$  structure is assigned as  $H_2C=C=CO^+$  on the grounds that this is the most stable  $H_2C_3O^+$  isomer, and it can be formed readily from fumaroyl dichloride during electron impact with only a 1,2-hydrogen shift, which is known to be facile in ions produced by electron bombardment.<sup>19</sup> The alternative  $HC=CH-CO^+$  isomer proposed in a theoretical study<sup>16</sup> cannot be excluded as a possibility although we had no evidence in our experiments with the  $H_2C_3O^+$  ion that more than one isomer was present. Of the  $H_3C_3O^+$  isomers investigated (see Table I), the most stable acyclic species by about 171 kJ mol<sup>-1</sup> has the protonated propadienone structure,  $H_2C=CH-CO^+$ . This structure can be produced from acrylic anhydride by cleavage of the  $H_2CCHC(O)-OC(O)CHCH_2$  bond without further rearrangement, and we thus expect the  $H_3C_3O^+$  ion in our study to be almost exclusively  $H_2C=CH-CO^+$ . The observation of linear semilogarithmic decays of these ions with added neutral reactants indicates that only one isomeric form of each ion is present in the flow tube.<sup>7,20</sup>

## Results and Discussion

**$H_nC_3O^+$  Chemistry.** A summary of the reaction chemistry of the  $H_nC_3O^+$  ions is given in Table II. It is evident, from the results presented in the table, that both of the ions  $HC_3O^+$  and  $H_2CCCO^+$  are relatively unreactive. Only the reactions with  $NH_3$  and  $CH_3CCH$  exhibit rate coefficients that are a substantial fraction of the collision rate. Also,  $HC_3O^+$  exhibits an association

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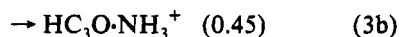
\* Abstract published in *Advance ACS Abstracts*, November 15, 1993.

**TABLE I: Thermochemistry of  $H_nC_3O^+$  ( $n = 0-3$ ),  $C_3O_2^+$ , and Related Neutrals (Values in Bold Type Denote Those Chosen for the Indicated Structures in This Work)**

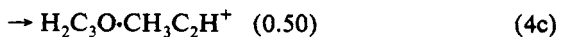
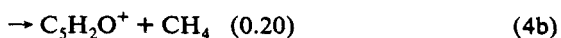
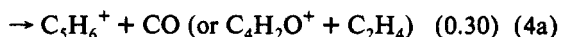
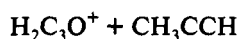
molecule/ion	$\Delta H_f^\circ$ (kJ mol <sup>-1</sup> )	method <sup>a</sup>	ref
$C_3O$			
$C=C=C=O^+$	>1252	R	this work
	1393	MNDO	11
$C=C=C=O$	<b>326</b>	PA	12
	282	MO	13
	390	MO	14
$HC_3O$			
$HC\equiv C-C=O^+$	<b>971</b>	AP	15
$HC=C=C=O$	< <b>299</b>	R	this work
$H_2C_3O$			
$H_2C=C=C=O^+$	<b>975</b>	IP	15
$HC=CH-CO^+$	996	MO <sup>b</sup>	16
$HC=CH-C=O^+$	1052	IP	15
	1089	MO <sup>b</sup>	16
$HC\equiv C-C-OH^+$	1120	MO <sup>b</sup>	16
$HC\equiv C-CHO^+$	1157	IP	17
	1216	MO <sup>b</sup>	16
$HC=CH-C=O$	138	est	15
$HC\equiv C-CHO$	115	est	15
$H_2C=C=C=O$	<b>95</b>	est	15
$H_3C_3O$			
$H_2C=CH-CO^+$	<b>751</b>	AP	15
	749	MNDO	18
$HC=CH-C=OH^+$	868	MO <sup>c</sup>	18
$HC\equiv C-CH-OH^+$	922	MO	18
$H_2C=C-COH^+$	969	MO <sup>c</sup>	18
$HC\equiv C-CH_2O^+$	950	AP	15
$H_2C=CHCO$	72		15
$C_3O_2$			
$O=C=C=C=O^+$	<b>929</b>	IP	15
	832	MNDO	11
$O=C=C=C=O$	-94		15

<sup>a</sup> AP = appearance potential determination, est = estimated value, IP = ionization potential determination, MNDO = MNDO calculation, MO = ab initio MO calculation; PA = ab initio calculation of proton affinity, and R = reaction study. <sup>b</sup> Referenced to  $\Delta H_f^\circ(H_2CCCO^+)$ . <sup>c</sup> Referenced to  $\Delta H_f^\circ(H_2CCHCO^+)$ .

product channel with  $NH_3$  while  $H_2C_3O^+$  similarly associates with  $CH_3CCH$ . In each case the association channel is competing with other bimolecular channels.



$$k_3 = 2.1 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$$



$$k_4 = 1.2 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$$

Although the association process (3b, 4c) is usually represented as the result of a bimolecular reaction, it is in fact a termolecular process, where the initial complex is stabilized by multiple collisions with the helium bath gas at 0.30 Torr. We have previously found association to be competitive with exothermic bimolecular channels in a number of systems,<sup>22,23</sup> and reactions 3b and 4c are two further examples. The association products

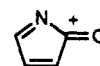
are strongly bound adducts in which covalent bonding occurs, but in the present cases the isomeric forms of the adducts are not known. In reaction 3b the most stable acyclic ion known is protonated pyruvonitrile,  $CH_3COCNH^+$ , but this product cannot be formed from  $HC_3O^+ + NH_3$  without substantial rearrangement. Also, examination of the variation of the adduct  $H_4C_3NO^+$  signal with  $NH_3$  flow showed that no rapid reaction with  $NH_3$  occurred. A rapid reaction would be expected if the  $H_4C_3NO^+$  product was  $CH_3COCNH^+$  in view of the fact that its proton affinity (PA) is known to be much less than that of  $NH_3$ . One possible product that has a PA greater than that of  $NH_3$  is the cyclic molecule protonated oxazole. Similarly, in reaction 4c the most stable  $H_6C_6O^+$  ion identified is the phenol cation  $C_6H_5OH^+$ .

A slightly slower association reaction was also observed to occur between  $H_2C_3O^+$  and HCN



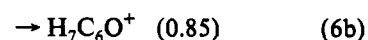
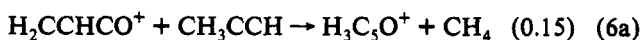
$$k_5 = 1.9 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$$

Again the isomeric form of the  $H_3C_4NO^+$  adduct is not known although one cyclic possibility could be the cation of pyrrol-2-one,



Alternatively, a six-membered aromatic molecular ion has also been suggested.<sup>5</sup> In reactions 3b, 4c, and 5 where the stabilized adduct is observed, some stability of the collision complex is inferred as it must survive sufficiently long for stabilization to occur via collisions with the helium bath gas.

The reaction chemistry of  $H_2CCHCO^+$  is also summarized in Table II. From this it is evident that this ion is even less reactive than either  $HC_3O^+$  or  $H_2C_3O^+$ . Again, association reactions, although occurring at low efficiency, are the major reaction channels, and the adduct ion  $H_7C_6O^+$  is the major product of the reaction between  $H_2CCHCO^+$  and  $CH_3CCH$



$$k_6 = 6.8 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$$

The most stable  $H_7C_6O^+$  species is protonated phenol,  $C_6H_5OH_2^+$ , which would correspond to a well depth of about 325 kJ mol<sup>-1</sup> below ground-state reactants on the  $H_7C_6O^+$  potential energy surface. Generally, when the complex ion formed as a result of association has a particularly stable configuration, then its lifetime with respect to unimolecular dissociation is long enough to allow stabilization via collisions with the bath gas to occur before the complex has time to dissociate back to reactants.<sup>24</sup> Consequently, at higher pressures the association process can move into the pressure-saturated regime and appear to be independent of pressure. It is the large density of states above a deep well that contributes to the longer lifetime. When the well depth between initial reactants and association product is shallower, corresponding to metastable isomeric forms, the average complex lifetime decreases with a consequent reduction in the association rate. Herbst et al.<sup>25</sup> predicted the presence of metastable isomers in several associating systems when statistical calculations overestimated the observed association rate coefficients. In view of the small association rate coefficient observed for the reaction between  $H_2CCHCO^+$  and  $CH_3CCH$  (reaction 6b) and the deep well to the potential product,  $C_6H_5OH_2^+$ , we

TABLE II: Reactions of H<sub>n</sub>C<sub>3</sub>O<sup>+</sup> + X<sup>p</sup>

X	HC <sub>3</sub> O <sup>+</sup>			H <sub>2</sub> CCCO <sup>+</sup>			H <sub>2</sub> CCHCO <sup>+</sup>				
	products	branching ratio	k <sub>obs</sub> <sup>a</sup>	products	branching ratio	k <sub>obs</sub> <sup>a</sup>	products	branching ratio	k <sub>obs</sub> <sup>a</sup>	k <sub>c</sub> <sup>a,b</sup>	
O <sub>2</sub>	no reaction		<0.001	CHO <sup>+</sup> + HCO + CO <sup>c</sup>	0.60	0.036	no reaction		<0.001	0.65	
				C <sub>2</sub> H <sub>2</sub> O <sup>+</sup> + CO <sub>2</sub> <sup>d</sup>	0.40						
NH <sub>3</sub>	CH <sub>3</sub> CO <sup>+</sup> + HCN <sup>e</sup>	0.55	0.21	NH <sub>4</sub> <sup>+</sup> + HC <sub>3</sub> O <sup>f</sup>	1.0	1.3	H <sub>3</sub> C <sub>3</sub> O-NH <sub>3</sub> <sup>+</sup>	1.0	0.021	2.2	
	HC <sub>3</sub> O-NH <sub>3</sub> <sup>+</sup>	0.45									
C <sub>2</sub> H <sub>2</sub>	no reaction		<0.001	C <sub>3</sub> O <sup>+</sup> + 2H <sub>2</sub> <sup>g</sup>	0.30	0.43	no reaction		<0.001	1.0	
				C <sub>3</sub> HO <sup>+</sup> + H <sub>2</sub> + H <sup>h</sup>	0.55						
				C <sub>3</sub> H <sub>2</sub> O <sup>+</sup> + H <sub>2</sub> <sup>i</sup>	0.15						
HCN	HC <sub>3</sub> O-HCN <sup>+</sup>	1.0	0.004	H <sub>2</sub> C <sub>3</sub> O-NCH <sup>+</sup>	1.0	0.19	C <sub>3</sub> H <sub>3</sub> O-HCN <sup>+</sup>	1.0	0.002	3.3	
CH <sub>3</sub> CCH	C <sub>3</sub> H <sub>5</sub> <sup>+</sup> + CO <sup>j</sup>	1.0	0.92	C <sub>3</sub> H <sub>6</sub> <sup>+</sup> + CO <sup>j</sup>	0.30		1.2	H <sub>3</sub> C <sub>3</sub> O <sup>+</sup> + CH <sub>4</sub> <sup>o</sup>	0.15	0.068	1.4
	(or C <sub>4</sub> HO <sup>+</sup> + C <sub>2</sub> H <sub>4</sub> ) <sup>k</sup>			(or C <sub>4</sub> H <sub>2</sub> O <sup>+</sup> + C <sub>2</sub> H <sub>4</sub> ) <sup>m</sup>				H <sub>3</sub> C <sub>3</sub> O-CH <sub>3</sub> C <sub>2</sub> H <sup>+</sup>	0.85		
				C <sub>3</sub> H <sub>2</sub> O <sup>+</sup> + CH <sub>4</sub> <sup>n</sup>	0.20						
				H <sub>2</sub> C <sub>3</sub> O-CH <sub>3</sub> C <sub>2</sub> H <sup>+</sup>	0.50						

<sup>a</sup> Rate coefficients in units of 10<sup>-9</sup> cm<sup>3</sup> s<sup>-1</sup>. <sup>b</sup> Calculated collision rate coefficients at 300 K using either Langevin theory when X has no dipole moment or the theory of Su and Chesnavich.<sup>21</sup> <sup>c</sup> ΔH<sup>o</sup> = -215 kJ mol<sup>-1</sup> for this channel. <sup>d</sup> ΔH<sup>o</sup> = -489 kJ mol<sup>-1</sup> for the H<sub>2</sub>CCO<sup>+</sup> isomer and -336 kJ mol<sup>-1</sup> for alternative possibility, HCCOH<sup>+</sup>. <sup>e</sup> Production of CH<sub>3</sub>CO<sup>+</sup> is exothermic by 137 kJ mol<sup>-1</sup> while production of CH<sub>2</sub>COH<sup>+</sup> is endothermic by 13 kJ mol<sup>-1</sup>. <sup>f</sup> Proton transfer is exothermic if ΔH<sub>f</sub><sup>o</sup> (HC<sub>3</sub>O) < 299 kJ mol<sup>-1</sup>. <sup>g</sup> Requires ΔH<sub>f</sub><sup>o</sup> (C<sub>3</sub>O<sup>+</sup>) < 1203 kJ mol<sup>-1</sup>. <sup>h</sup> Requires ΔH<sub>f</sub><sup>o</sup> (C<sub>3</sub>HO<sup>+</sup>) < 985 kJ mol<sup>-1</sup>. <sup>i</sup> Requires ΔH<sub>f</sub><sup>o</sup> (C<sub>3</sub>H<sub>2</sub>O<sup>+</sup>) < 1203 kJ mol<sup>-1</sup>. <sup>j</sup> Formation of several possible C<sub>3</sub>H<sub>5</sub><sup>+</sup> isomers is energetically possible. <sup>k</sup> If C<sub>4</sub>HO<sup>+</sup> + C<sub>2</sub>H<sub>4</sub> are the products, then ΔH<sub>f</sub><sup>o</sup> (C<sub>4</sub>HO<sup>+</sup>) < 1105 kJ mol<sup>-1</sup>. <sup>l</sup> Calculated for ionized cyclopentadiene, c-C<sub>3</sub>H<sub>6</sub><sup>+</sup>, but other C<sub>3</sub>H<sub>6</sub><sup>+</sup> isomers may also be formed. <sup>m</sup> Requires ΔH<sub>f</sub><sup>o</sup> (C<sub>4</sub>H<sub>2</sub>O<sup>+</sup>) < 1109 kJ mol<sup>-1</sup>. <sup>n</sup> Exothermic by more than 33 kJ mol<sup>-1</sup> using a value of ΔH<sub>f</sub><sup>o</sup> (C<sub>3</sub>H<sub>2</sub>O<sup>+</sup>) < 1203 kJ mol<sup>-1</sup>. <sup>o</sup> Requires ΔH<sub>f</sub><sup>o</sup> (H<sub>3</sub>C<sub>3</sub>O<sup>+</sup>) < 1012 kJ mol<sup>-1</sup>. <sup>p</sup> Note: reactions with several other neutrals were investigated but were observed as "no reactions". These experiments gave the following upper limits for k<sub>obs</sub> values, with all three H<sub>n</sub>C<sub>3</sub>O<sup>+</sup> ions, as follows: H<sub>2</sub>, 0.0005; CO, N<sub>2</sub>, and CH<sub>4</sub>, 0.001; H<sub>2</sub>O, 0.01.

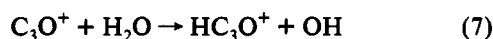
TABLE III: Reactions of C<sub>3</sub>O<sup>+</sup> + X

X	products	branching ratio	k <sub>obs</sub> <sup>a</sup>	k <sub>c</sub> <sup>a,b</sup>	-ΔH <sup>o</sup> (kJ mol <sup>-1</sup> ) <sup>c</sup>
H <sub>2</sub>	HC <sub>3</sub> O <sup>+</sup> + H	1.0	0.8	1.5	>63
CH <sub>4</sub>	HC <sub>3</sub> O <sup>+</sup> + CH <sub>3</sub>	1.0	0.83	1.1	>60
NH <sub>3</sub>	HC <sub>3</sub> O <sup>+</sup> + NH <sub>2</sub> <sup>d</sup>	}	1.9	2.2	>46
	C <sub>3</sub> O-NH <sub>3</sub> <sup>+</sup>				
H <sub>2</sub> O	HC <sub>3</sub> O <sup>+</sup> + OH	1.0	0.5	2.4	>0
C <sub>2</sub> H <sub>2</sub>	C <sub>4</sub> H <sub>2</sub> <sup>+</sup> + CO	0.45	1.08	1.0	>69
	C <sub>3</sub> O-C <sub>2</sub> H <sub>2</sub> <sup>+</sup>	0.55			
HCN	HC <sub>3</sub> N <sup>+</sup> + CO <sup>d</sup>	}	2.9	3.3	>24
	C <sub>3</sub> O-CN <sup>+</sup> + H				
CO	C <sub>4</sub> O <sub>2</sub> <sup>+</sup>	1.0	0.07	0.79	
N <sub>2</sub>	no reaction		<0.001	0.73	
O <sub>2</sub>	CO <sup>+</sup> + 2CO	0.20	0.22	0.67	>232
	C <sub>3</sub> O <sub>2</sub> <sup>+</sup> + O	0.80			
CO <sub>2</sub>	no reaction		<0.001	0.78	>74

<sup>a</sup> Rate coefficients in units of 10<sup>-9</sup> cm<sup>3</sup> s<sup>-1</sup>. <sup>b</sup> Calculated collision rate coefficients at 300 K using either Langevin theory when X has no dipole moment or the theory of Su and Chesnavich.<sup>21</sup> <sup>c</sup> Enthalpies are derived from the values tabulated in Table I and from the compilation of Lias et al.<sup>15</sup> <sup>d</sup> Product ratio not determined. See text for discussion.

feel that protonated phenol is unlikely as the association product in this case and that a less stable isomer of H<sub>7</sub>C<sub>6</sub>O<sup>+</sup> is produced.

**C<sub>3</sub>O<sup>+</sup> and C<sub>3</sub>O<sub>2</sub><sup>+</sup> Chemistry.** A summary of the reaction chemistry of C<sub>3</sub>O<sup>+</sup> is given in Table III. The reaction enthalpies shown assume a value for ΔH<sub>f</sub><sup>o</sup>(C<sub>3</sub>O<sup>+</sup>) > 1252 kJ mol<sup>-1</sup> which is based on the observation of the rapid reaction

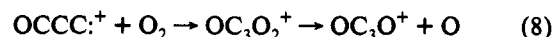


$$k_7 = 5 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$$

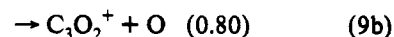
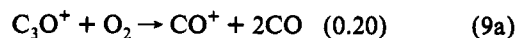
Terlouw et al. have determined ΔH<sub>f</sub><sup>o</sup>(HC<sub>3</sub>O<sup>+</sup>) = 971 ± 5 kJ mol<sup>-1</sup> from the measured appearance potential of HC<sub>3</sub>O<sup>+</sup>.<sup>26</sup> This value, when linked with Botschwina's ab initio calculation for the proton affinity of C<sub>3</sub>O of 885 kJ mol<sup>-1</sup>,<sup>12</sup> yields ΔH<sub>f</sub><sup>o</sup>(C<sub>3</sub>O) = 326 kJ mol<sup>-1</sup>. Accurate branching ratios for the product channels of the reactions with NH<sub>3</sub> and HCN could not be found under the conditions of the experiment because of low signal intensities and the fact that, in each case, one of the product ions was adjacent in mass to the reactant C<sub>3</sub>O<sup>+</sup> ion.

There is a marked increase in the reactivity of C<sub>3</sub>O<sup>+</sup> ions over the reactions presented in Table II for H<sub>n</sub>C<sub>3</sub>O<sup>+</sup>, and this can be attributed to the carbene character of O=C=C=C<sup>+</sup>. McElvany

et al. have attributed the reactivity of C<sub>3</sub>O<sup>+</sup> toward O<sub>2</sub> as being due to the carbene character of C<sub>3</sub>O<sup>+</sup> such that complexation of O<sub>2</sub> occurs at the carbene end being followed by O atom loss.<sup>11</sup>



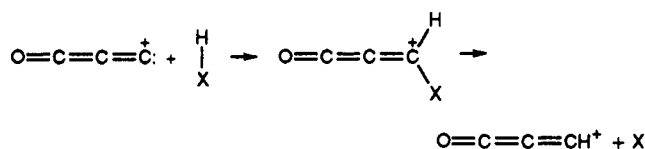
In their low-pressure study (≤ 1 × 10<sup>-6</sup> Torr) they reported just one channel (reaction 8) with a rate coefficient of 1.3 × 10<sup>-10</sup> cm<sup>3</sup> s<sup>-1</sup>. In the present work, at the higher pressures of the flow tube, we have observed two channels



$$k_9 = 2.2 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$$

The additional channel 9a possibly arises from partial collisional deactivation of the C<sub>3</sub>O-O<sub>2</sub><sup>+</sup> complex, increasing its lifetime and thus allowing a further exit channel from the complex to occur.

Much of the reaction chemistry of C<sub>3</sub>O<sup>+</sup> in this study can also be interpreted in terms of carbene insertion into the H-X σ bond as



in the reactions with H<sub>2</sub>, CH<sub>4</sub>, NH<sub>3</sub>, and H<sub>2</sub>O. This type of process did not occur in the reactions with C<sub>2</sub>H<sub>2</sub> and HCN where H atom transfer is not as energetically favorable and may be endothermic.

The reaction which controls the chemistry of C<sub>3</sub>O<sup>+</sup> in interstellar clouds is the rapid reaction with H<sub>2</sub>



$$k_{10} = 8.0 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$$

and this reaction ensures that C<sub>3</sub>O<sup>+</sup> will have a relatively short lifetime in dense clouds.

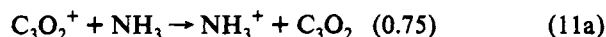
A summary of the reaction chemistry of C<sub>3</sub>O<sub>2</sub><sup>+</sup> is presented in Table IV. It is evident that C<sub>3</sub>O<sub>2</sub><sup>+</sup> exhibits less reactivity

TABLE IV: Reactions of  $C_3O_2^+ + X$ 

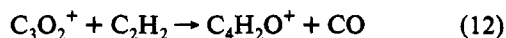
X	products	branching ratio	$k_{obs}^a$	$k_c^{a,b}$	$-\Delta H^\circ$ (kJ mol <sup>-1</sup> )
H <sub>2</sub>	no reaction		<0.0001	1.5	
NH <sub>3</sub>	NH <sub>3</sub> <sup>+</sup> + C <sub>3</sub> O <sub>2</sub>	0.75	2.0	2.1	43
	H <sub>2</sub> C <sub>2</sub> O <sup>+</sup> + HNCO	0.25			
CH <sub>4</sub>	C <sub>3</sub> O <sub>2</sub> H <sup>+</sup> + CH <sub>3</sub>	1.0	0.3	1.05	109 <sup>c</sup>
H <sub>2</sub> O	C <sub>3</sub> O <sub>2</sub> H <sub>2</sub> O <sup>+</sup>	1.0	0.001	2.0	<i>d</i>
C <sub>2</sub> H <sub>2</sub>	C <sub>4</sub> H <sub>2</sub> O <sup>+</sup> + CO	1.0	1.2	1.0	<i>e</i>
HCN	C <sub>3</sub> O <sub>2</sub> H <sub>2</sub> CN <sup>+</sup>	1.0	0.015	3.2	
N <sub>2</sub>	no reaction		<0.001	0.70	
CO	no reaction		<0.001	0.76	
O <sub>2</sub>	CO <sup>+</sup> + CO + CO <sub>2</sub>	0.10	0.12	0.63	194
	CO <sub>2</sub> <sup>+</sup> + 2CO	0.75			
	C <sub>2</sub> O <sub>2</sub> <sup>+</sup> + CO <sub>2</sub>	0.15			
CO <sub>2</sub>	no reaction		<0.001	0.74	288 <sup>f</sup>

<sup>a</sup> Rate coefficients in units of 10<sup>-9</sup> cm<sup>3</sup> s<sup>-1</sup>. <sup>b</sup> Calculated collision rate coefficients at 300 K using either Langevin theory when X has no dipole moment or the theory of Su and Chesnavich.<sup>21</sup> <sup>c</sup> Calculated for the products H<sub>2</sub>C=CO<sup>+</sup> + HNCO. <sup>d</sup> The observation of this channel yields  $\Delta H_f^\circ(C_3O_2H^+) < 709$  kJ mol<sup>-1</sup>; i.e.,  $PA(C_3O_2) > 727$  kJ mol<sup>-1</sup>. <sup>e</sup> The observation of this channel yields  $\Delta H_f^\circ(C_4H_2O^+) < 1268$  kJ mol<sup>-1</sup>. <sup>f</sup> Calculated using  $D(OC-CO^+) = 93.7$  kJ mol<sup>-1</sup> (ref 27).

than does C<sub>3</sub>O<sup>+</sup> as only the reactions with NH<sub>3</sub> and C<sub>2</sub>H<sub>2</sub> proceed at the collision rate.



$$k_{11} = 2.0 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$$



$$k_{12} = 1.2 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$$

The extent of rearrangement required in both reactions 11b and 12 supports mechanisms whereby the reactions proceed through tight collision complexes.

### Conclusion

The ion chemistry of H<sub>n</sub>C<sub>3</sub>O<sup>+</sup> ions is characterized by a noticeable lack of reactivity with the major known neutral constituents of interstellar clouds. As a consequence, all the H<sub>n</sub>C<sub>3</sub>O<sup>+</sup> ions, viz HC<sub>3</sub>O<sup>+</sup>, H<sub>2</sub>C<sub>3</sub>O<sup>+</sup>, and H<sub>2</sub>CCHCO<sup>+</sup>, are likely

to have relatively long cloud lifetimes with their major loss processes being dissociative recombinations. Association reactions of H<sub>n</sub>C<sub>3</sub>O<sup>+</sup> were found to present an important outcome for several reactions, and these can lead to the efficient formation of new and more complex molecular ions. The much greater reactivity of the C<sub>3</sub>O<sup>+</sup> ion over the other ions investigated in this study can be attributed to the carbene character of C<sub>3</sub>O<sup>+</sup>.

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