

**DETECTION OF THE  $N_3$  FREE RADICAL BY LASER MAGNETIC RESONANCE AT 6.08  $\mu\text{m}$** 

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Lines in the  $\nu_3$  (antisymmetric stretch) fundamental band of the  $N_3$  radical in the  $\tilde{X}^2\Pi_g$  state have been detected by CO laser magnetic resonance. The observations have been assigned to P, Q, and R lines in the fundamental vibration-rotation band and lead to a precise determination of the vibrational interval, the ground-state rotational constant and its dependence on the antisymmetric stretching vibration:  $\nu_3 = 1644.67903(24) \text{ cm}^{-1}$ ,  $B_0 = 0.431438(16) \text{ cm}^{-1}$  and  $\alpha_3 = 0.004357(18) \text{ cm}^{-1}$ .

**1. Introduction**

As Walsh pointed out in his memorable series of papers [1], triatomic molecules of general formula  $BAC$  with 16 or fewer valence electrons are expected to be linear in their ground and low-lying states. Several of the possible 15-electron molecules that can be formed from first-row elements have now been identified and characterised. All conform well to Walsh's predictions. The best studied in this group are  $BO_2$  [2],  $CO_2^+$  [3],  $NCO$  [4] and  $N_2O^+$  [5]. The molecule  $N_3$  is another example, but there has been comparatively little work so far to determine its structure. It was first identified by Thrush [6], who recorded its electronic spectrum at 270 nm by flash photolysis of hydrazoic acid. The same spectrum was recorded later at higher resolution by Douglas and Jones [7]. They analysed the (0, 0) band of a  $^2\Sigma_u^+ - \tilde{X}^2\Pi_g$  transition and made the first determination of rotational and spin-orbit parameters for  $N_3$ . More recently, the photoelectron spectrum of  $N_3$  has been recorded by Dyke et al. [8], but as usual, this provides information on the positive ion  $N_3^+$  rather than the parent molecule. Nevertheless, the study did show that  $N_3$  can be formed efficiently by a chemical reaction, namely that of F atoms with  $HN_3$ .

Direct structural information is hard to come by in the case of  $N_3$ . It has no permanent dipole moment in its ground vibrational level (although its equilibrium structure may well be asymmetrical [9]) and consequently no rotational spectrum. However,

it should have two strong infrared transitions involving the bending vibration  $\nu_2$  (probably in the region of  $500 \text{ cm}^{-1}$ ) and the antisymmetric stretch  $\nu_3$ . Unfortunately, none of the previous studies involved  $\nu_3$  vibrational intervals and there was considerable uncertainty as to its value. We embarked on a search for the  $\nu_3$  band several years ago using carbon monoxide laser magnetic resonance (LMR), confident that the band would fall somewhere within the range of the laser ( $1250\text{--}2030 \text{ cm}^{-1}$  [10]). The search concentrated on the region near  $1900 \text{ cm}^{-1}$ , guided by the analogous transition in  $NCO$  [11]. These searches were not successful. However, the  $\nu_3$  band of  $N_3$  has been detected very recently by FTIR techniques around  $1645 \text{ cm}^{-1}$  [12]. On searching in this region, we quickly found strong signals. This result highlights the weakness of LMR as a search technique. It is undoubtedly much more tedious to search over tens or even hundreds of CO laser lines than it is to integrate over a broad bandwidth with an interferometer. Unless the search region is fairly well defined by other means, either theoretical or experimental, the LMR spectroscopist is very likely to give up before finding the spectrum.

This paper reports the observation and analysis of preliminary measurements in the (1, 0) band of the  $N_3$  radical by carbon monoxide LMR.

## 2. Experimental details and observations

The LMR spectrometer has been described elsewhere [13]. Briefly, it involves a flowing CO laser cooled with liquid nitrogen. A diffraction grating is used to select the laser line and the laser is stabilised by locking to the top of the gain curve. The absorption cell is located inside the laser cavity. Molecular transitions are tuned into coincidence with the laser frequency by application of an external magnetic field which can be varied between 0 and 1.6 T. Zeeman modulation at 5 kHz is applied with a pair of cooled Helmholtz coils, capable of generating a field of up to 12 mT (peak-to-peak). The magnetic flux densities are measured with a proton NMR gaussmeter.

The  $N_3$  radical was generated in the absorption cell by the reaction of F atoms with  $HN_3$ , as described by Dyke et al. [8]. Hydrazoic acid was generated in situ by heating a mixture of 10 g of stearic acid with 0.65 g of sodium azide to 90°C. The pressure of  $HN_3$  under these conditions was about 1 Torr (133 Pa). F atoms were generated by passing 5%  $F_2$  in He through a microwave discharge. The total pressure in the reaction cell was about 0.5 Torr (66 Pa).

Signals associated with the  $\nu_3$  band of  $N_3$  were detected with four different laser lines. Only spectra in perpendicular polarization have been investigated so far. The spectrum for the  $R_1(3/2)$  transition using the  $P(16)_{18,17}$  CO laser line at  $1647.05403\text{ cm}^{-1}$  is shown in fig. 1. The details of the observed resonances are given in table 1. The laser wavenumbers have been calculated from the Dunham coefficients of Koch, Vanek and Evenson, who have recently measured the frequencies of several CO laser lines directly using heterodyne techniques [14]. It was found to be comparatively easy to produce saturation (or Lamb) dips on most of the lines. Indeed for the  $R_1(5/2)$  transition recorded with the  $P(9)_{19,18}$  line at  $1647.81217\text{ cm}^{-1}$ , almost all the resonances occur at very low fields and are not resolvable with Doppler linewidths. The measurements given in this case all correspond to Lamb dips. It was also possible to resolve  $^{14}\text{N}$  hyperfine structure on some lines in the  $R_1(3/2)$  spectrum.

One curious characteristic of our observations is that, under certain circumstances, we observed the same spectra in emission rather than absorption. We

have not been able to characterise the necessary conditions fully but shall investigate these effects further.

## 3. Analysis and discussion

The assignments of the LMR spectra were made by comparison of the observed Zeeman patterns with the predictions of a computer program, using the parameters given by Douglas and Jones [7]. The process proved to be quite straightforward. For example the  $R_1(3/2)$  spectrum shown in fig. 1 shows the classic Zeeman structure associated with this transition. Details are given in table 1.

As usual with this type of spectrum, our observations have been confined to  $N_3$  in the  $^2\Pi_{3/2}(F_1)$  spin component. The measurements depend primarily on the band origin  $\nu_3$  (or more strictly the separation between the  $^2\Pi_{3/2}$  components) and on the rotational constants for the two vibrational levels involved. There is also a very weak dependence on other parameters for a  $^2\Pi$  state such as  $A$  (spin-orbit

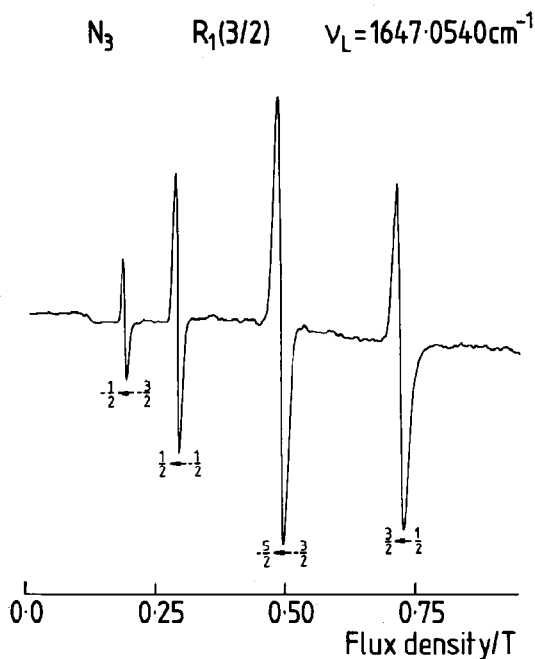


Fig. 1. The LMR spectrum associated with the  $R_1(3/2)$  transition of the  $\nu_3$  fundamental band of the  $N_3$  radical, observed with the  $P(16)_{18,17}$  CO laser line at  $1647.05403\text{ cm}^{-1}$  in perpendicular polarisation. The  $M_J$  assignments are given below each transition.

Table 1

Observed resonances in the laser magnetic resonance spectrum of the  $\nu_3$  band of the  $N_3$  radical

| Transition | $M_J$ |      | Laser frequency<br>( $\text{cm}^{-1}$ ) | Flux density<br>(mT) | $10^4(\text{obs} - \text{calc})$<br>( $\text{cm}^{-1}$ ) | $\partial\nu/\partial B_0$<br>(MHz/G) |
|------------|-------|------|-----------------------------------------|----------------------|----------------------------------------------------------|---------------------------------------|
| $R_1(5/2)$ | 1/2   | -1/2 | 1647.81217 <sup>a)</sup>                | 42.5 <sup>b)</sup>   | 0.6                                                      | 0.387 <sup>c)</sup>                   |
|            | -7/2  | -5/2 | 1647.81217                              | 66.1 <sup>b)</sup>   | 0.4                                                      | 0.251                                 |
|            | 3/2   | 1/2  | 1647.81217                              | 93.1 <sup>b)</sup>   | 1.2                                                      | 0.174                                 |
|            | -5/2  | -3/2 | 1647.81217                              | 444.0 <sup>b)</sup>  | -2.1                                                     | 0.043                                 |
| $R_1(3/2)$ | -1/2  | -3/2 | 1647.05403                              | 154.3                | -4.9                                                     | 1.484                                 |
|            | 1/2   | -1/2 | 1647.05403                              | 259.2                | 2.7                                                      | 0.917                                 |
|            | -5/2  | -3/2 | 1647.05403                              | 473.3                | 0.6                                                      | 0.486                                 |
|            | 3/2   | 1/2  | 1647.05403                              | 716.9                | 1.6                                                      | 0.462                                 |
| $Q_1(3/2)$ | 1/2   | 3/2  | 1644.28899                              | 1430.0               | 0.8                                                      | -1.210                                |
|            | -1/2  | 1/2  | 1644.28899                              | 1569.0               | -0.8                                                     | -1.008                                |
| $P_1(9/2)$ | -7/2  | -5/2 | 1640.73237                              | 1090.0               | 0.5                                                      | -0.647                                |
|            | -5/2  | -3/2 | 1640.73237                              | 1352.0               | -0.5                                                     | -0.522                                |

<sup>a)</sup> Carbon monoxide laser wavenumber, determined from the Dunham coefficients of ref. [14].<sup>b)</sup> Observed as a Lamp dip.<sup>c)</sup> The values quoted for the transition tuning rates are calculated from the parameter values given in table 3.

coupling),  $\gamma$  (spin-rotation coupling) and  $D$  (centrifugal distortion). To obtain the best estimates of these quantities, we have refitted the optical data of Douglas and Jones as a  $^2\Sigma^+ - ^2\Pi$  transition, using the  $N^2$  formulation of Brown et al. [15] for the  $^2\Pi$  state. All lines which were marked by the authors as being assigned to two or more transitions were excluded from the fit. The parameter values obtained are given in table 2. The values for  $A''$  and  $B''$  for the  $X^2\Pi_g$  state agree reasonably well with those given by Douglas and Jones ( $-72.26$  and  $0.43117 \text{ cm}^{-1}$ , respectively).

We have used the measurements from the LMR spectrum in table 1 to determine four parameters for  $N_3$  in the  $\tilde{X}^2\Pi_g$  state:  $\nu_3$ ,  $B_0$ ,  $\alpha_3$  and  $g_L$ . The Hund's case (a) basis set was truncated at  $\Delta J=3$  in the cal-

culations and the data were assigned equal weights in the fit. The standard deviation of the fit was  $2.3 \times 10^{-4} \text{ cm}^{-1}$ . The residuals obtained are given in table 1 and the corresponding set of molecular parameters in table 3. The values obtained for the pa-

Table 3

Molecular parameters for the  $N_3$  radical in its  $\tilde{X}^2\Pi_g$  state determined from observations of the  $\nu_3$  band by LMR spectroscopy

| Parameter <sup>a)</sup> | Present work                 | FTIR study<br>[12]     |
|-------------------------|------------------------------|------------------------|
| $\nu_3$                 | 1644.67903(24) <sup>b)</sup> | 1644.6784              |
| $A_0$                   | -71.2675 <sup>c,d)</sup>     | -71.2722 <sup>d)</sup> |
| $\alpha_4$              | 0.3810 <sup>e)</sup>         | -0.3810                |
| $B_0$                   | 0.431438(16)                 | 0.43145                |
| $10^2\alpha_3$          | 0.4357(18)                   | 0.4386                 |
| $10^6D_0$               | 0.2091 <sup>c)</sup>         | 0.1886                 |
| $10^2\gamma_0$          | -0.227 <sup>c,d)</sup>       | -0.074 <sup>d)</sup>   |
| $g_L$                   | 0.98512(86)                  | -                      |

<sup>a)</sup> Value in  $\text{cm}^{-1}$ , where appropriate.<sup>b)</sup> The figures in parentheses represent one standard deviation of the least-squares fit, in units of the last quoted decimal place.<sup>c)</sup> Parameter constrained to this value in the fit, obtained from a fit of the optical data of Douglas and Jones [7].<sup>d)</sup> Value for effective parameter, obtained with  $A_D$  constrained to zero.<sup>e)</sup> Parameter constrained to this value in the fit, taken from the FTIR results of Brazier et al. [12].

Table 2

Molecular parameters (in  $\text{cm}^{-1}$ ) for  $N_3$  obtained from a refit of the (0,0) band of the  $^2\Sigma_u^+ - \tilde{X}^2\Pi_g$  transition [7]

|           |                            |                |              |
|-----------|----------------------------|----------------|--------------|
| $\nu_0$   | 36738.751(3) <sup>a)</sup> | $A''$          | -71.2675(45) |
| $B'$      | 0.432453(61)               | $10^2\gamma''$ | -0.227(55)   |
| $10^6D'$  | 0.203(30)                  | $10^3p''$      | 0.95(26)     |
| $B''$     | 0.431261(61)               | $10^4q''$      | -41(15)      |
| $10^6D''$ | 0.209(30)                  |                |              |

<sup>a)</sup> The figures in parentheses represent one standard deviation of the least-squares fit, in units of the last quoted decimal place.

rameters in the FTIR study are also given in table 3 for comparison. There is excellent agreement between the two sets of parameters.

The primary quantity determined by these observations is the antisymmetric stretching frequency  $\nu_3$  ( $1644.6790\text{ cm}^{-1}$ ). The value obtained is rather low by comparison with the corresponding frequency for NCO ( $1920.6\text{ cm}^{-1}$  [11]) but is high when compared with the values for  $\text{BO}_2$  ( $1278.26\text{ cm}^{-1}$  [16]) and for  $\text{CO}_2^+$  ( $1423.08\text{ cm}^{-1}$  [17]). There does not seem to be any simple rationalisation of this set of numbers. The other parameter values given in table 3 are more or less as expected. In particular, the value for the orbital  $g$  factor  $g_L$  of 0.9851 is significantly less than unity. As has been discussed elsewhere [11,18], this arises in part from Renner-Teller mixing of the bending vibrational levels of the  $^2\Pi_g$  state and in part from vibronic mixing of excited  $^2\Sigma_u$  and  $^2\Delta_u$  states. When the Renner-Teller effect has been better characterised for  $\text{N}_3$ , it will be possible to separate these two contributions.

Much further work remains to be done on the LMR spectrum of  $\text{N}_3$ . In addition to characterising the  $^{14}\text{N}$  hyperfine interaction from the measurement and analysis of the Lamb dip signals, we hope to be able to detect hot bands involving the bending vibration to learn more about the Renner-Teller coupling. We shall also study the conditions for the observation of the signals in emission in more detail.

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might well have been a long time before we looked in the right wavenumber region to observe the spectrum. We also thank Dave Robinson, who was responsible for the construction of the LMR spectrometer and was involved in the early searches for  $\text{N}_3$ , and Jonathan Towle, who provided valuable back-up services for the present study.

### References

- [1] A.D. Walsh, J. Chem. Soc. (1953) 2266.
- [2] J.W.C. Johns, Can. J. Phys. 39 (1961) 1738.
- [3] D. Gauyacq, C. Larcher and J. Rostas, Can. J. Phys. 57 (1979) 1634.
- [4] R.N. Dixon, Phil. Trans. Roy. Soc. A 252 (1960) 165.
- [5] J.H. Callomon and F. Creutzberg, Phil. Trans. Roy. Soc. A 277 (1974) 159.
- [6] B.A. Thrush, Proc. Roy. Soc. A 235 (1956) 143.
- [7] A.E. Douglas and W.E. Jones, Can. J. Phys. 43 (1965) 2216.
- [8] J.M. Dyke, N.B.H. Jonathan, A.E. Lewis and A. Morris, Mol. Phys. 47 (1982) 1231.
- [9] T.W. Archibald and J.R. Sabin, J. Chem. Phys. 55 (1971) 1821.
- [10] W. Rohrbeck, A. Hinz, P. Nelle, M.A. Gondal and W. Urban, Appl. Phys. B 31 (1983) 139.
- [11] C.E. Barnes, J.M. Brown, A.D. Fackerell and T.J. Sears, J. Mol. Spectry. 92 (1982) 485.
- [12] C.R. Brazier, P.F. Bernath, J.B. Burkholder and C.J. Howard, J. Chem. Phys., to be published.
- [13] W. Seebass, J. Werner, W. Urban, E.R. Comben and J.M. Brown, Mol. Phys. 62 (1987) 161.
- [14] M. Koch, M. Vanek and K.M. Evenson, to be published.
- [15] J.M. Brown, E.A. Colbourn, J.K.G. Watson and F.D. Wayne, J. Mol. Spectry. 74 (1979) 294.
- [16] K. Kawaguchi, E. Hirota and C. Yamada, Mol. Phys. 44 (1981) 509.
- [17] K. Kawaguchi, C. Yamada and E. Hirota, J. Chem. Phys. 82 (1985) 1174.
- [18] A. Carrington, A.R. Fabris, B.J. Howard and N.J.D. Lucas, Mol. Phys. 20 (1971) 961.